

Gravitational waves and BMS symmetries

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2. Symmetries and conservation laws in GW

3. BMS Symmetries

A brief intro on GW discovery

In 1918, Einstein predicted gravitational waves (GW)

154 Gesamtsitzung vom 14. Februar 1918. - Mitteilung vom 31. Januar

Über Gravitationswellen.

Von A. EINSTEIN.

(Vorgelegt am 31. Januar 1918 [s. oben S. 79].)

In Sep. 2015, first detection of GW by LIGO (after > 20 years of work)



Detection of gravitational waves

LIGO-Virgo collaboration





Gravity in the strong regieme

- Einstein equation: Nonlinear system of PDE
- Describing interaction between matter and spacetime
- This can be broken (perturbatively) into two sub-problems

1. GW generation Given a source, find the gravitational field it produces

- Post-Newtonian formalism: $v/c \ll 1$
- Gravitational self force & Extreme mass ratio inspirals: $\ m/M \ll 1$
- Numerical Relativity
- 2. Back-reaction of gravitational field on the source

Use conservation laws to study radiation reaction

Symmetries and conservation laws in GW

Asymptotic Symmetries



- Far from the source, the spacetime is asymptotically flat
- Poincaré symmetries are asymptotically restored
- One can apply Noether's theorem and associate charges to the system

$$\mathcal{E}, \quad \mathcal{P}_i, \quad \mathcal{J}_i, \quad \mathcal{K}_i$$

• These charges are not conserved, but they obey flux-balance laws:

$$\frac{dQ}{dt} = -\mathcal{F}_Q$$

Why balance equations? [Peters, Mathews '63]

- Newtonian gravity $\mathcal{O}(v/c)^0$ No radiation. Conservation laws enough to solve Kepler problem
- First PN approximation Balance equations replace conservation laws due to radiation

$$\frac{dE_{mech}}{dt} = -\mathcal{F}_E$$

$$E_{mech} = -\frac{Gm_1m_2}{2r}, \qquad \mathcal{F}_E = \frac{G}{5c^5} \, \overleftarrow{I}_{ij} \, \overleftarrow{I}_{ij}, \qquad I_{ij} = \int d^3x \rho \, x_i \, x_j$$

• Using these in balance equation determines the orbit [Pictures: Maggiore]



Poincare flux balance laws

Einstein quadrupole formula [Einstein '18]

$$\dot{\mathcal{E}} = -\frac{1}{5} \frac{G}{c^5} \widetilde{I}_{ij} \widetilde{I}_{ij} + \mathcal{O}(1/c^7)$$

Angular momentum flux [Epstein, Wagoner '75, Thorne '80]

$$\dot{\mathcal{J}}_i = -\frac{2}{5} \frac{G}{c^5} \epsilon_{ijk} \ddot{I}_{jl} \ddot{I}_{kl} + \mathcal{O}(1/c^7)$$

Linear momentum flux [Thorne '80]

$$\dot{\mathcal{P}}_{i} = -\frac{2}{63} \frac{G}{c^{7}} \stackrel{(4)}{I} \stackrel{(3)}{_{ijk}} \stackrel{(3)}{I}_{jk} + \mathcal{O}(1/c^{9})$$

Center of mass flux [Kozameh '16, Blanchet, Faye '18]

$$\dot{\mathcal{G}}_{i} = \mathcal{P}_{i} - \frac{1}{21} \frac{G}{c^{7}} \left(\stackrel{(3)}{I}_{jk} \stackrel{(3)}{I}_{ijk} - \stackrel{(2)}{I}_{jk} \stackrel{(4)}{I}_{ijk} \right) + \mathcal{O}(1/c^{9})$$

Our task: Find exact equations and extend the list

BMS Symmetries

- Found originally in the 60's and recently extended [Barnich, Compère, Campiglia, etc.]
- An infinite dimensional extension of the Poincaré group
- Relations found with Weinberg soft theorems and memory effects [Strominger et.al]

Bondi asymptotic expansion

- Setup: Asymptotic expansion of metric in 1/r
- Metric in Bondi gauge [retarded coordinates (u, r, θ^A)]

$$\begin{split} ds^2 &= -c^2 \, du^2 - 2c \, dudr + r^2 \gamma_{AB} \, d\theta^A d\theta^B \\ &+ \frac{2G}{c^2 r} m \, du^2 + \frac{1}{r} \frac{4G}{3c^2} N_A \, dud\theta^A + r \, \underline{C_{AB}} \, d\theta^A d\theta^B \\ &+ \dots \end{split}$$

- Bondi data: $m = m(u, \theta^A), N_A = N_A(u, \theta^A)$ Bondi mass, angular momentum aspect, $C_{AB}(u, \theta^A)$ Bondi shear
- Einstein equations

$$\partial_u m = -\frac{c^3}{8G}\dot{C}_{AB}\dot{C}^{AB} + \frac{c^4}{4G}D_A D_B\dot{C}^{AB},$$

• The coordinate transformation $x \to x + \xi$ is a symmetry if

$$\xi = \underbrace{T(\theta)\partial_u}_{\text{supertranslation}} + \underbrace{\epsilon^{BC}\partial_B\Phi(\theta)\partial_C}_{\text{superrotation}} + \underbrace{\gamma^{BC}\partial_B\Psi(\theta)\partial_C}_{\text{superboost}}$$

• Called BMS symmetry. It contains the Poincaré algebra

	Т	Φ	Ψ
$Y_{0,0}$	time translation	-	-
$Y_{1,m}$	spatial translation	rotations	boosts

• Corresponding BMS charges

$$\begin{split} \mathcal{P}_{T} &\equiv \frac{1}{c} \oint_{S} T \, \boldsymbol{m} \,, & Supermomentum \\ \mathcal{J}_{\Phi} &\equiv \frac{1}{2} \oint_{S} \epsilon^{AB} \partial_{B} \Phi N_{A} , & S\text{-angular momentum} \\ \mathcal{K}_{\Psi} &\equiv \frac{1}{2c} \oint_{S} \gamma^{AB} \partial_{B} \Psi N_{A} & S\text{-center of mass} \end{split}$$

• These charges are NOT conserved but obey flux balance equations

$$\dot{P}_T = \frac{1}{c} \oint_S T \dot{m}$$
$$= \frac{1}{c} \oint_S T \left(-\frac{c^3}{8G} \dot{C}_{AB} \dot{C}^{AB} + \frac{c^4}{4G} D_A D_B \dot{C}^{AB} \right)$$

• Similarly for superLorentz charges

Multipole expansion

- Why multipole expansion? PN hierarchy of fluxes
- Multipole expansion of radiation field $(n_i = \frac{x_i}{r}, N_L = n_{i_1} \cdots n_{i_\ell})$

$$C_{ij} = \sum_{\ell=2}^{+\infty} a_\ell \left(N_{L-2} \ \boldsymbol{U}_{ijL-2} - \frac{b_\ell}{c} N_{aL-2} \ \epsilon_{ab(i} \boldsymbol{V}_{j)bL-2} \right)^{TT}$$

in terms of mass and spin radiative multipole moments.

• Relation to source multipoles (from PN formalism)

$$\boldsymbol{U}_{L} = \overset{(\ell)}{\boldsymbol{I}}_{L} + \mathcal{O}\left(\frac{G}{c^{3}}\right), \qquad \boldsymbol{V}_{L} = \overset{(\ell)}{\boldsymbol{J}}_{L} + \mathcal{O}\left(\frac{G}{c^{3}}\right)$$

Poincarè flux balance equations

• Supermomentum balance equation

$$\dot{P}_T = \frac{1}{c} \oint_S T \left(-\frac{c^3}{8G} \dot{C}_{AB} \dot{C}^{AB} + \frac{c^4}{4G} D_A D_B \dot{C}^{AB} \right)$$

• Energy T = 1

$$\dot{\mathcal{E}} = -\sum_{\ell=2}^{+\infty} rac{G}{c^{2\ell+1}} \mu_\ell igg\{ \dot{U}_L \dot{U}_L + rac{b_\ell b_\ell}{c^2} \dot{V}_L \dot{V}_L igg\}$$

Leading terms

$$\dot{\mathcal{E}} = -\frac{G}{c^5} \left(\frac{1}{5} I_{ij}^{(3)} I_{ij}^{(3)}\right) - \frac{G}{c^7} \left(\frac{1}{189} I_{ijk}^{(4)} I_{ijk}^{(4)} + \frac{16}{45} J_{ij}^{(3)} J_{ij}^{(3)}\right) + O(c^{-9})$$

- Perfect match with the literature [Thorne '80, Blanchet, Faye '18]
- Similar results for linear momentum and center of mass [Kozameh '16]

• Proper BMS balance laws [Compère, Oliveri, AS '19, see also Nichols '18]

$$T(\theta^A) = T_L N_L \qquad \ell \ge 2$$

- New equations with the same PN order as the energy and angular momentum
- Example: Octupolar super angular momentum

$$\dot{\mathcal{J}}_{ijk} - \frac{2}{7c^2} \, u \, \dot{\boldsymbol{V}}_{ijk} = -\frac{G}{c^5} \left(\frac{6}{35} \, \epsilon_{pq\langle i} \, \dot{\boldsymbol{U}}_{j|p|} \boldsymbol{U}_{k\rangle q} \right) + \mathcal{O}(c^{-7})$$

Summary

- New flux balance equations
- In principle relevant for radiation reaction forces
- specifies gravitational memory effect

Outlook

- Wrtie BMS charges in terms of source variables
- New differential equations for source parameters

Thank you very much