# 5D SUSY Black Ring/CFTs Higher Derivative Terms 

Hesam Soltanpanahi

University of the Witwatersrand Johannesburg, South Africa

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## \& 5 D SUGRA $+\mathrm{H} . \mathrm{D}$. C.

Compactification of $M$-theory on a $\mathrm{CY}_{3}$ results in $\mathcal{N}=2$ SUGRA in 5D.
The bosonic action up to 4th order in superconformal formalism is

$$
I=\frac{1}{16 \pi G_{5}} \int d^{5} x \sqrt{|g|}\left(\mathcal{L}_{0}+\mathcal{L}_{1}\right)
$$

where

$$
\begin{aligned}
\mathcal{L}_{0} & =\partial_{a} \mathcal{A}_{\alpha}^{i} \partial^{a} \mathcal{A}_{i}^{\alpha}+\left(2 \nu+\mathcal{A}^{2}\right) \frac{D}{4}+\left(2 \nu-3 \mathcal{A}^{2}\right) \frac{R}{8}+\left(6 \nu-\mathcal{A}^{2}\right) \frac{v^{2}}{2}+2 \nu_{I} F_{a b}^{I} v^{a b} \\
& +\frac{1}{4} \nu_{I J}\left(F_{a b}^{I} F^{J a b}+2 \partial_{a} X^{I} \partial^{a} X^{J}\right)+\frac{e^{-1}}{24} C_{I J K} \epsilon^{a b c d e} A_{a}^{I} F_{b c}^{J} F_{d e}^{K} \\
\mathcal{L}_{1} & =\frac{c_{2 I}}{24}\left(\frac{1}{16 e} \epsilon_{a b c d e} A^{I a} R^{b c f g} R^{d e}{ }_{f g}+\frac{1}{8} X^{I} C^{a b c d} C_{a b c d}+\frac{1}{12} X^{I} D^{2}+\frac{1}{6} F^{I a b} v_{a b} D\right. \\
& +\frac{1}{3} X^{I} C_{a b c d} v^{a b} v^{c d}+\frac{1}{2} F^{I a b} C_{a b c d} v^{c d}+\frac{8}{3} X^{I} v_{a b} \widehat{\mathcal{D}}^{b} \widehat{\mathcal{D}}_{c} v^{a c} \\
& +\frac{4}{3} X^{I} \hat{\mathcal{D}}^{a} v^{b c} \widehat{\mathcal{D}}_{a} v_{b c}+\frac{4}{3} X^{I} \widehat{\mathcal{D}}^{a} v^{b c} \widehat{\mathcal{D}}_{b} v_{c a}-\frac{2}{3 e} X^{I} \epsilon_{a b c d e} v^{a b} v^{c d} \widehat{\mathcal{D}}_{f} v^{e f} \\
& +\frac{2}{3 e} F^{I a b} \epsilon_{a b c d e} v^{c f} \widehat{\mathcal{D}}_{f} v^{d e}+e^{-1} F^{I a b} \epsilon_{a b c d e} v^{c}{ }_{f} \widehat{\mathcal{D}}^{d} v^{e f} \\
& \left.-\frac{4}{3} F^{I a b} v_{a c} v^{c d} v_{d b}-\frac{1}{3} F^{I a b} v_{a b} v^{2}+4 X^{I} v_{a b} v^{b c} v_{c d} v^{d a}-X^{I}\left(v^{2}\right)^{2}\right)
\end{aligned}
$$

related to the mixed gauge-gravitational CS term by SUSY trensformations.
$C_{I J K}$ : Intersection numbers of internal space $\mathrm{CY}_{3}$
$C_{2} I$ : Second Chern class of internal space $\mathrm{CY}_{3}$

$$
\begin{gathered}
\mathcal{A}^{2}=\mathcal{A}_{\alpha}^{i} \mathcal{A}_{i}^{\alpha}, \quad v^{2}=v_{a b} v^{a b}, \\
\nu=\frac{1}{6} C_{I J K} X^{I} X^{J} X^{K}, \quad \nu_{I}=\frac{1}{2} C_{I J K} X^{J} X^{K}, \quad \nu_{I J}=C_{I J K} X^{K} .
\end{gathered}
$$

Weyl multiplet $\rightarrow$ the metric, a 2-form auxiliary field, $v_{a b}$, a scalar auxiliary field $D$, a gravitino $\psi_{\mu}^{i}$ and an auxiliary Majorana spinor $\chi^{i}$

Vector mutiplet $\rightarrow 1$-form gauge field $A^{I}$, a scalar auxiliary field $X^{I}$ and a gaugino $\Omega^{I i}$ (where $I=1, \cdots, n_{v}$ count the number of vector multiplets) and $i=1,2$ is an $S U(2)$ doublet index.

Hyper multiplet $\rightarrow$ auxiliary scalar fields $\mathcal{A}_{\alpha}^{i}$ and a hyperino $\zeta^{\alpha} \alpha=1, \cdots, 2 r$ refers to $\operatorname{USP}(2 r)$ group.

The near horizon of SUSY black ring is given by

$$
\begin{aligned}
& d s^{2}=l_{A d S^{2}}^{2}\left(-r^{2} d t^{2}+\frac{d r^{2}}{r^{2}}\right)+l_{S^{1}}^{2}\left(d \psi+e_{0} r d t\right)^{2}+l_{S^{2}}^{2}\left(d \theta^{2}+\sin ^{2} \theta d \phi^{2}\right), \\
& A^{I}=e^{I} r d t-\frac{p^{I}}{2} \cos \theta d \phi+a^{I}\left(e_{0} r d t+d \psi\right), \quad X^{I}=\frac{p^{I}}{l_{A d S^{2}}}, \quad D=\frac{12}{l_{A d S 2^{2}}^{2}}, \\
& Q^{I}=-4 C_{I J K} p^{J} a^{K}, \quad e^{I}+e_{0} a^{I}=0, \quad v_{\theta \phi}=\frac{3}{8} l_{A d S^{2}} \sin \theta,
\end{aligned}
$$

$\theta \& \phi$ : The coordinates of a usual 2-sphere.
$\psi$ : The coordinate of ring, $\psi \sim \psi+4 \pi$.
The radii are given by the magnetic charges $p^{I}$,

$$
l_{A d S_{2}}=l_{S^{2}}=e_{0} l_{S^{1}}=\frac{1}{2}\left(\frac{1}{6} C_{I J K} p^{I} p^{J} p^{K}+\frac{1}{12} c_{2 I} p^{I}\right)^{1 / 3}
$$

The macroscopic entropy is

$$
S_{\text {mac }}=\frac{2 \pi}{e_{0}}\left(\frac{1}{6} C_{I J K} p^{I} p^{J} p^{K}+\frac{1}{6} c_{2 I} p^{I}\right)=2 \pi \sqrt{\frac{\hat{q}_{0}\left(C_{I J K} p^{I} p^{J} p^{K}+c_{2 I} p^{I}\right)}{6}}
$$

in which

$$
\widehat{q}_{0}=\frac{1}{e_{0}{ }^{2}}\left(\frac{1}{6} C_{I J K} p^{I} p^{J} p^{K}+\frac{1}{6} c_{2 I} p^{I}\right)
$$

\& c -extremization approach
$\diamond$ Near Horizon Isometry
The isometry of the metric are $S L(2, R) \times U(1) \times S O(3)$ generated by

$$
\begin{array}{rlrl}
K_{1} & =t \partial_{t}-r \partial_{r}, & & K_{2}=\frac{1}{2}\left(t^{2}+r^{-2}\right) \partial_{t}-r t \partial_{r}-\frac{e_{0}}{r} \partial_{\psi} \\
K_{3} & =\partial_{t}, & K_{4}=\partial_{\psi}, \\
J^{3} & =-i \partial_{\phi}, & J^{ \pm}=e^{ \pm i \phi}\left(-i \partial_{\theta} \pm \cot \theta \partial_{\phi}\right) .
\end{array}
$$

The first two parts of the near horizon metric is locally $\mathrm{AdS}_{3}$ and this permit us to use the c-extremization approach to find the associated central charge.
[Kraus and Larsen (05)]
The first step is choosing an appropriate ansatz,

$$
\begin{aligned}
d s^{2} & =l_{A d S_{3}}^{2} d s_{A d S_{3}}^{2}+l_{S^{2}}^{2} d s_{S^{2}}^{2} \\
A^{I} & =e^{I} r d t-\frac{p^{I}}{2} \cos \theta d \phi+a^{I}\left(e_{0} r d t+d \psi\right)
\end{aligned}
$$

Then by extremizing thec-function

$$
c=6 l_{A}{ }^{3} l_{S}{ }^{2}\left(\mathcal{L}_{0}+\mathcal{L}_{1}\right)
$$

with respect to the $l_{A}$ and $l_{S}$ we find their values in terms of the magnetic charges and the value of c-function at these radii gives the average of left and right central charges.

By doing these calculation one finds,

$$
\begin{aligned}
& l_{A d S_{3}}=2 l_{S^{2}}=\left(\frac{1}{6} C_{I J K} p^{I} p^{J} p^{K}+\frac{1}{12} c_{2 I} p^{I}\right)^{1 / 3} \\
& A^{I}=e^{I} r d t-\frac{p^{I}}{2} \cos \theta d \phi+a^{I}\left(e_{0} r d t+d \psi\right), \quad X^{I}=\frac{p^{I}}{l_{A d S^{2}}}, \quad D=\frac{12}{l_{A d S^{2}}^{2}}, \\
& Q^{I}=-4 C_{I J K} p^{J} a^{K}, \quad e^{I}+e_{0} a^{I}=0, \quad v_{\theta \phi}=\frac{3}{8} l_{A d S^{2}} \sin \theta,
\end{aligned}
$$

and the value of c-function at this extremum point is given by

$$
\left.c\right|_{e x t .}=\frac{1}{2}\left(c_{L}+c_{R}\right)=C_{I J K} p^{I} p^{J} p^{K}+\frac{3}{4} c_{2 I} p^{I}
$$

There is a precise agreement between the above solution and the results of entropy function formalism.

For the associated dual CFT the gravitational anomaly yields to the difference between left and right central charges,

$$
c_{L}-c_{R}=\frac{1}{2} c_{2 I} p^{I}
$$

Thus the left and right central charges are given by

$$
c_{L}=C_{I J K} p^{I} p^{J} p^{K}+c_{2 I} p^{I}, \quad c_{R}=C_{I J K} p^{I} p^{J} p^{K}+\frac{1}{2} c_{2 I} p^{I}
$$

Using the Cardy formula (in extremal limit) to compute the microscopic entropy

$$
S_{\mathrm{mic}}^{(c)}=2 \pi \sqrt{\frac{c_{L} \widehat{q}_{0}}{6}}=\frac{2 \pi}{e_{0}}\left(\frac{1}{6} C_{I J K} p^{I} p^{J} p^{K}+\frac{1}{6} c_{2 I} p^{I}\right)=S_{m a c}
$$

Note that this entropy associated to a chiral CFT based on the $S L(2, R)$ part of the isometry of near horizon geometry of black ring such that $L_{0}$, $L_{1}$ and $L_{-1}$ are respectively proportional to $K_{1}, K_{2}$ and $K_{3}$ isometries of the geometry.

Thus there is a Virasoro algebra with the above central charge

$$
\left[L_{m}^{(c)}, L_{n}^{(c)}\right]=(m-n) L_{m+n}^{(c)}+\frac{c_{L}}{12}\left(m^{3}-m\right) \delta_{m+n}
$$

such that

$$
\begin{aligned}
& t \partial_{t}-r \partial_{r} \\
& \partial_{t} \Rightarrow L_{0}^{(c)}, L_{ \pm 1}^{(c)} \\
& \frac{1}{2}\left(t^{2}+r^{-2}\right) \partial_{t}-r t \partial_{r}-\frac{e_{0}}{r} \partial_{\psi}
\end{aligned}
$$

and

$$
c_{L}=C_{I J K} p^{I} p^{J} p^{K}+c_{2 I} p^{I}
$$

\& Kerr/CFT approach
$\diamond$ Asymptotic Symmetry
Near horizon metric of 5D Black Ring:

$$
\begin{gathered}
d s^{2}=\frac{\tilde{p}^{2}}{4}\left(-r^{2} d t^{2}+\frac{d r^{2}}{r^{2}}\right)+\frac{\tilde{p}^{2}}{4}\left(d \theta^{2}+\sin ^{2} \theta d \phi^{2}\right)+\tilde{L}^{2}\left(d \psi+e_{0} r d t\right)^{2} \\
\tilde{p}^{3} \equiv \frac{1}{6} C_{I J K} p^{I} p^{J} p^{K}+\frac{1}{12} c_{2 I} p^{I} \quad \tilde{L} \equiv \frac{\tilde{p}}{2 e_{0}}
\end{gathered}
$$

Consistent boundary conditions ( $t, r, \theta, \phi, \psi$ ):

$$
h_{\mu \nu} \sim \mathcal{O}\left(\begin{array}{ccccc}
r^{2} & 1 / r^{2} & 1 / r & r & 1 \\
& 1 / r^{3} & 1 / r^{2} & 1 / r^{2} & 1 / r \\
& & 1 / r & 1 / r & 1 / r \\
& & & 1 / r & 1 \\
& & & & 1
\end{array}\right)
$$

The generators associated to these boundary conditions are given by

$$
\zeta_{n}=-e^{-i n \psi} \partial_{\psi}-i n r e^{-i n \psi} \partial_{r}
$$

such that satisfy a Virasoro algebra

$$
i\left[\zeta_{m}, \zeta_{n}\right]=(m-n) \zeta_{m+n}
$$

$\diamond$ Computing the central charge
1st Point: To extend the Kerr/CFT approach for theory with higher derivative corrections it is useful to do the calculations in non-basis coordinates.

[Azeyanagi, Compere, Ogawa, Tachikawa, and Terashima(09)]

The vielbeins associated to near horizon geometry of black ring are

$$
\begin{array}{ll}
e^{\hat{t}}=\frac{\tilde{p}}{2} r d t, \quad e^{\widehat{r}}=\frac{\tilde{p}}{2 r} d r, \quad e^{\widehat{\theta}}=\frac{\tilde{p}}{2} d \theta \\
e^{\hat{\phi}}=\frac{\tilde{p}}{2} \sin \theta d \phi, \quad e^{\hat{\psi}}=\tilde{L}\left(d \psi+e_{0} r d t\right)
\end{array}
$$

and the variation of the metric under Virasoro generators is given by

$$
\begin{array}{ll}
\mathcal{L}_{\zeta_{n}} e^{\hat{t}}=i n e^{-i n \psi} e^{\hat{t}}, & \mathcal{L}_{\zeta_{n}} e^{\widehat{r}}=-e_{0} n^{2} e^{-i n \psi}\left(e^{\hat{\psi}}-e^{\hat{t}}\right) \\
\mathcal{L}_{\zeta_{n}} e^{\hat{\theta}}=\mathcal{L}_{\zeta_{n}} e^{\hat{\phi}}=0, & \mathcal{L}_{\zeta_{n}} e^{\hat{\psi}}=i n e^{-i n \psi}\left(e^{\hat{\psi}}-2 e^{\hat{t}}\right)
\end{array}
$$

2nd Point: The Kerr/CFT approach was extended to the case with CS term and it was shown that for a theory with gravity and also other fields, the central charge is not affected by non-gravitational fields.
[Compere, Murata, and Nishioka(09)]

Using the BBC method the central charge is given by
[ Barnich and Brandt (01), Barnich and Compere (07), Compere (07)]

$$
c^{(k)}=\left.12 i \int_{\partial \Sigma} \mathbf{k}_{\zeta_{n}}^{i n v}\left[\mathcal{L}_{\zeta_{-n}} g ; g\right]\right|_{n^{3}}
$$

$I_{n^{3}}$ : Stands for the term of order $n^{3}$
$\partial \Sigma$ : Spatial boundry

$$
\begin{aligned}
\mathbf{k}_{\zeta_{n}}^{i n v}\left[\mathcal{L}_{\zeta_{-n}} g ; g\right]= & -2\left[\mathbf{X}_{c d} \mathcal{L}_{\zeta_{n}} \nabla^{c} \zeta_{-n}^{d}+\left(\mathcal{L}_{\zeta_{n}} \mathbf{X}\right)_{c d} \nabla^{[c} \zeta_{-n}^{d]}+\mathcal{L}_{\zeta_{n}} \mathbf{W}_{c} \zeta_{-n}^{c}\right] \\
& -\mathbb{E}\left[\mathcal{L}_{\zeta_{n}} g, \mathcal{L}_{\zeta_{-n}} g ; g\right]
\end{aligned}
$$

Covariant derivatives: are defined with respect to the original metric $g$. $\mathbf{X}$ and $\mathbf{W}$ and $E$ : are given by

$$
\begin{array}{r}
\left(\mathbf{W}^{c}\right)_{c_{3} c_{4} c_{5}}=-2 \nabla_{d} Z^{a b c d} \epsilon_{a b c_{3} c_{4} c_{5}}=2\left(\nabla_{d} \mathbf{x}^{c d}\right)_{c_{3} c_{4} c_{5}} \\
\mathbf{E}_{c_{3} c_{4} c_{5}}= \\
\frac{1}{2}\left(-\frac{3}{2} Z^{a b c d} \delta g_{c}^{e} \wedge \delta g_{e d}+2 Z^{a c d e} \delta g_{c d} \wedge \delta g_{e}^{b}\right) \epsilon_{a b c_{3} c_{4} c_{5}}
\end{array}
$$

in which

$$
Z^{a b c d}=\frac{\delta^{c o v} L}{\delta R_{a b c d}}
$$

[ Azeyanagi, Compere, Ogawa, Tachikawa, and Terashima(09)]

The important points for calculating the central charge are

1- The Lie derivative of vielbeins with respect to the diffeomorphisms
2- The isometry $S L(2, R) \times U(1)$
3- The $t-\psi$ reflection symmetry of the near horizon geometry

Doing some algebric calculations the central charge associated to the Virasoro algebra is derived

$$
c^{(k)}=-12 e_{0} \int_{\Sigma} Z_{a b c d} \epsilon^{a b} \epsilon^{c d} \operatorname{vol}(\Sigma)=\frac{6 e_{0}}{\pi} S_{\mathrm{mac}}
$$

In the last equality we used the Iyer-Wald formula for macroscopic entropy of a black hole which is generalization of Bekenstein-Hawking formula when the higher derivative terms are appeared.

The cental charge is

$$
c^{(k)}=C_{I J K} p^{I} p^{J} p^{K}+c_{2 I} p^{I}=c_{L}
$$

Note that this central charge equals to the left central charge computed by c-extremization formalism.

This equality was shown for SUSY black ring without higher derivative corrections.

$$
c^{(k)}=c_{L}=c_{R}=C_{I J K} p^{I} p^{J} p^{K}
$$

and

$$
S_{\mathrm{mic}}^{(k)}=S_{\mathrm{mic}}^{(c)}=S_{\mathrm{mac}}
$$

As in the other application of Kerr/CFT approach, to compute the microscopic entropy we should derive the Frolov-Thorne temperature. The temperature is an intrinsic feature of metric and its definition is not corrected by higher derivative terms.
[Azeyanagi, Compere, Ogawa, Tachikawa and Terashima (09)]
So as usual one can find the Frolov-Thorne temperature from the $t \psi$ cross term of near horizon geometry

$$
T_{\mathrm{FT}}=\frac{1}{\pi e_{0}}
$$

Often there is a factor 2 in the denominator but remember that the period of $\psi$ coordinate in our notation is $4 \pi$.

The microscopic entropy of supersymmetric black ring in Kerr/CFT approach can be computed by other form of Cardy formula,

$$
S_{\mathrm{mic}}^{(k)}=\frac{\pi^{2}}{3} c^{(k)} T_{F T}=\frac{2 \pi}{e_{0}}\left(\frac{1}{6} C_{I J K} p^{I} p^{J} p^{K}+\frac{1}{6} c_{2 I} p^{I}\right)=S_{\mathrm{mic}}^{(c)}=S_{\mathrm{mac}}
$$

As we expect this microscopic entropy equals to the microscopic entropy calculated by c-extremization formalism and also equals to the macroscopic entropy.

\% Summary:


Thank you for your attention

